# **CHIANTI**

An Astrophysical Database for Emission Line Spectroscopy

#### CHIANTI TECHNICAL REPORT No. 15

Computing level populations

### 1 Overview

The calculation of level populations of an ion is perhaps the single most important part of CHIANTI. This document describes the software and methods used to perform the calculation within IDL.

Most users are recommended to use the routine ch\_pops (Sect. 3) for calculating the populations.

## 2 The pop\_solver routine

The workhorse of the CHIANTI IDL software is pop\_solver, which takes the atomic data, creates the population process array, and then performs the matrix inversion to yield the level populations.

There is a two-step process in calling pop\_solver. For example,

```
IDL> input=ch_setup_ion('o_6')
IDL> pop_solver, input, 5e5, 1e9, pop
```

The first call loads the atomic data for the ion O VI into the structure input. The call to pop\_solver then uses the atomic data to compute the populations (pop) at the specified temperature  $(5 \times 10^5 \text{ K})$  and electron number density  $(10^9 \text{ cm}^{-3})$ .

The output pop is simply an array of populations for all of the ion's levels. If you need information on the identification of the levels, then you are recommended to use the routine ch\_pops instead (Section 3).

The routine ch\_setup\_ion has a number of different options which can switch off certain processes, or modify data. Section 4 summarizes these options.

As of CHIANTI 9, new routines were added that take the atomic data from input and reformat them into rate matrices that form the matrix solved by pop\_solver. More details of these "load rates" routines are given in Appendix B.

## 3 The ch\_pops routine

A more user-friendly way of obtaining the level populations is to use the routine ch\_pops.pro. For example:

```
IDL> pop=ch_pops('fe_13')
```

This prints the populations of the most populous levels to the screen, while information for all levels is output to the pop structure. The tags of the output are:

DENS	DOUBLE	1.0000000e+10
TEMP	DOUBLE	1778279.4
LEVEL	STRUCT	-> <anonymous> Array[749]</anonymous>
RADTEMP	FLOAT	-1.00000
RPHOT	FLOAT	-1.00000

```
PROTON STRING 'yes'

VERSION STRING 'CHIANTI 8.0.2'

DATE STRING 'Wed Jun 28 15:36:26 2017'

SUM_MWL INT 0

SUM_MWL_COEFFS FLOAT -1.00000
```

The tag level is a structure array, with the tags:

INDEX	INT	1
TERM	STRING	'3s2 3p2 3PO'
POP	DOUBLE	0.11817674

The level population is given in the tag pop and it is given as the population relative to the population of the ion as a whole, so summing the pop values over all levels gives 1.

By default, populations are calculated at an electron number density,  $N_{\rm e}$ , of  $10^{10}~{\rm cm^{-3}}$  and the  $T_{\rm max}$  of the ion (computed with ch\_tmax) is used for the temperature. Keywords exist to switch off proton rates (/noprot) and level-resolved ionization and recombination rates (/noionrec) where these are available for the ion. Photon excitation and non-Maxwellian distributions can be included using the standard CHIANTI keywords (see CHIANTI User Guide).

Note that ch\_pops is a wrapper for the older routine show\_pops, which is itself a wrapper for calling pop\_solver. ch\_pops was written to standardize the input notation with other CHIANTI routines.

# 4 Atomic data setup (ch\_setup\_ion)

The default call to ch\_setup\_ion results in all of the ion's data files being loaded into the output structure. The core data-sets of energy levels, A-values, and electron collision strengths are always loaded. The secondary data-sets of proton rates and level-resolved ionization and recombination rates are loaded as long as the data files exist for the ions. They can be switched off using the /NOPROT and /NOIONREC keywords, respectively.

Photon excitation (and stimulated emission) do not have atomic data files since the rates depend on the A-values. They are switched on by the user by specifying RPHOT, the distance from the emitting source center in source radius units. The blackbody radiation temperature is set with the input RADTEMP.

For proton rates, pop\_solver needs to know the proton density. This is computed self-consistently from the electron temperature (assumed to be the same as the proton temperature), and ion fraction file and the element abundance file. The optional inputs IONEQ\_FILE and ABUND\_FILE are used to specify these files. If not set, then the default files !IONEQ\_FILE and !ABUND\_FILE are used.

The simple call given in Sect. 2 will be sufficient for most users, but if you are calling ch\_setup\_ion from within another code and want to give users full access to the options, then the call would be:

## A Document history

Version 0.4, 5-Mar-2019. Added Sect. B.

### B The "load rates" routines

With CHIANTI 9, a significant update to the IDL software was made. Two routines that were added are ch\_load\_ion\_rates and ch\_load\_2ion\_rates. These routines essentially take the atomic data that was read by ch\_setup\_ion and then reformat them into atomic rates that enter into the matrix equation that is solved by pop\_solver. Previously this part of the code was contained in pop\_solver itself. Now pop\_solver calls out to the load rate routines.

Generally users will not need to run the load rates routines, but here we summarize how they work.

For most ions, the procedure is simply:

```
IDL> t=[1e5,2e5,3e5]
IDL> rates=ch_load_ion_rates('o_4',t)
```

where "t" is the temperature array for which the rates are computed. The result is a structure with the following tags:

N_LEVELS	LONG	204
AA	DOUBLE	Array[204, 204]
AAX	DOUBLE	Array[204, 204]
PPR	DOUBLE	Array[3, 204, 204]
QQ	DOUBLE	Array[3, 204, 204]
TEMP	FLOAT	Array[3]
ION_DATA	STRUCT	-> <anonymous> Array[1]</anonymous>
MULT	FLOAT	Array[204]
SUM_MWL_COEFFS	DOUBLE	Array[3]
SUMTST	INT	0

The rate matrices are: radiative decay rates (AA), photoexcitation and stimulated emission rates (AAX), proton rate coefficients (PPR), and electron rate coefficients (QQ). The ION\_DATA structure is simply the atomic data structure returned by ch\_setup\_ion.

If pop\_solver sees that the tag ion\_data.autostr exists, then it means that the ion has autoionization data (i.e., the .auto file exists). This requires a special two-ion atomic model whereby the CHIANTI model for the next ionization stage is added to the ion model. An example of how this is done for O VI is as follows:

```
IDL> rates1=ch_load_ion_rates('o_6',t)
IDL> rates2=ch_load_ion_rates('o_7',t)
IDL> rates=ch_load_2ion_rates(rates1,rates2)
```

The structure rates is then used for creating the atomic rates matrix used by pop\_solver. Note that this structure has an expanded set of tags:

N_LEVELS	LONG	972
AA	DOUBLE	Array[972, 972]
QQ	DOUBLE	Array[3, 972, 972]
AAX	DOUBLE	Array[972, 972]
PPR	DOUBLE	Array[3, 972, 972]
IONIZ	DOUBLE	Array[3, 972, 972]
RR	DOUBLE	Array[3, 972, 972]
AI	DOUBLE	Array[972, 972]
DC	DOUBLE	Array[3, 972, 972]
DR	DOUBLE	Array[3, 972, 972]

In particular: level-resolved ionization rate coefficients (IONIZ), level-resolved radiative recombination rate coefficients (RR), autoionization rates (AI), dielectronic capture rate coefficients (DC), and level-resolved dielectronic recombination rate coefficients (DR). As of CHIANTI 9, the IONIZ and DR matrices only contain data for transitions between the two ground states.

If you compare the rates1 and rates structures then you will see they have 923 and 972 levels, respectively (as of CHIANTI 9). This is because the 49 levels of the O VII CHIANTI model have been added to the O VI model. The level populations are calculated by pop\_solver for all 972 levels, but the array is then truncated to 923 levels so that only the O VI populations are returned. If you would like to see the populations for all 972 levels, then use the /all\_levels keyword input to pop\_solver.